

Λύμαν άλφα
Λαβύρινθος

Deciphering Cosmic Reionization History Encrypted In the Galaxy UV Continua

Hiroya Umeda (Utokyo)

Based on

1. *Umeda et al. (2024), ApJ, 971, 124*
2. *Umeda et al. (2025a), ApJS, 277, 37*
3. *Umeda et al. (2025b), on arXiv*

Ευχαριστούμε
τους διοργανωτές!

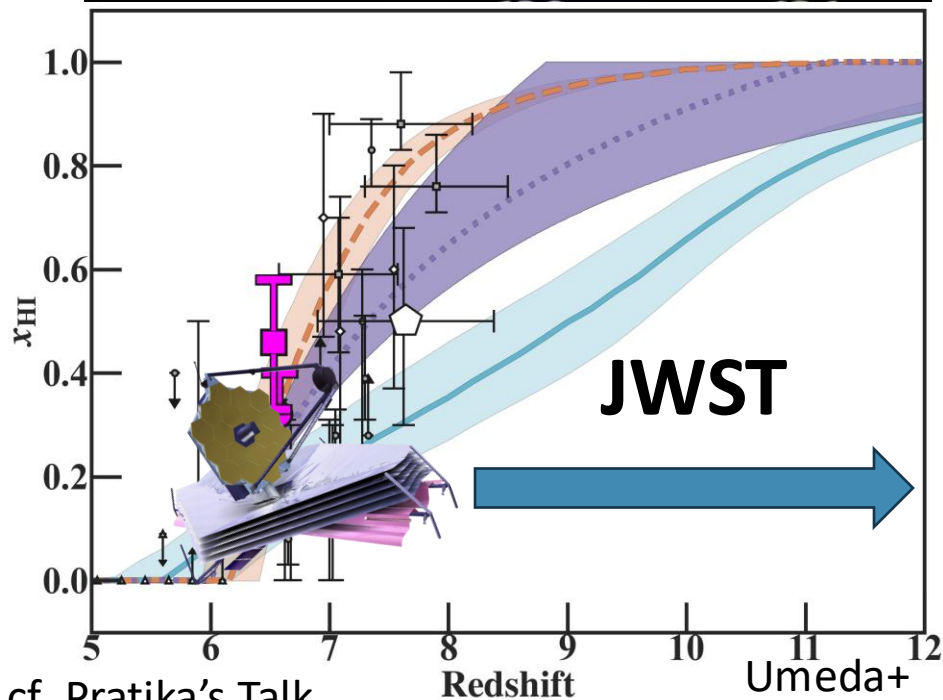
Wavelength (micron)

Escaping the Lyman Labyrinth

JWST Dives Into the Neutral Universe

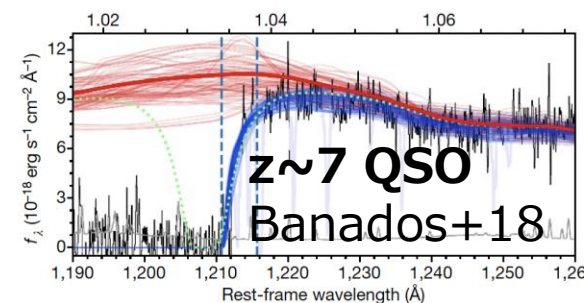
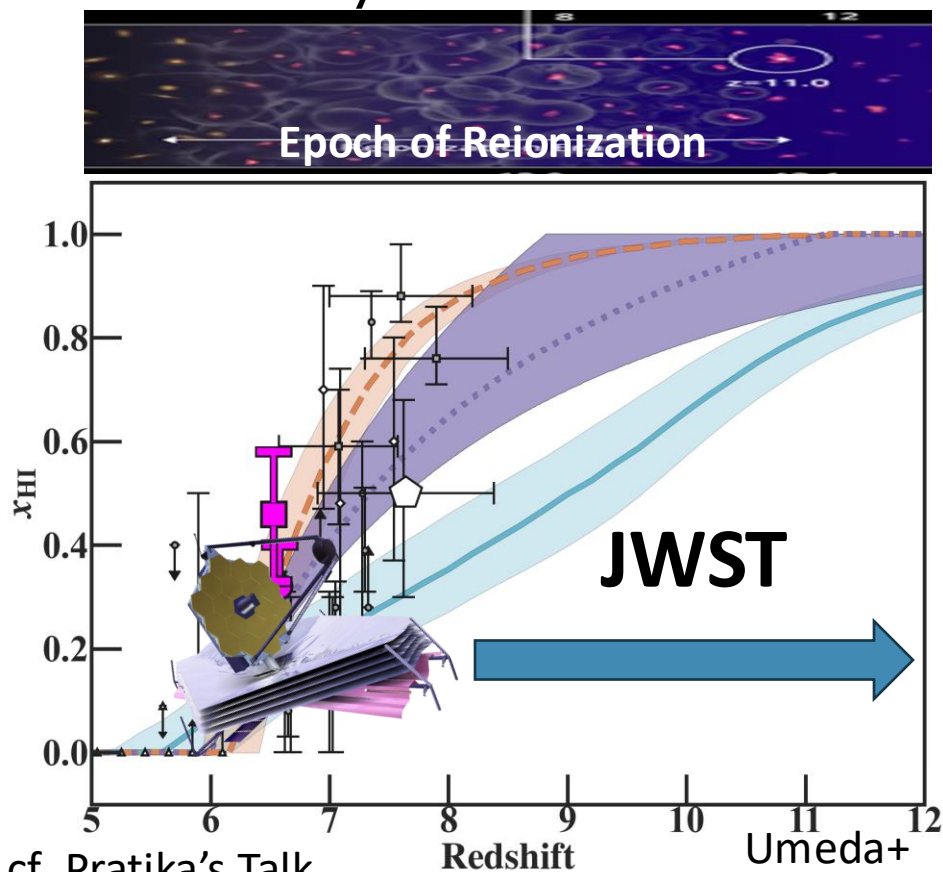
Nature of IGM via Galaxy Observation

Ly α Emission and ... Continuum Absorption?



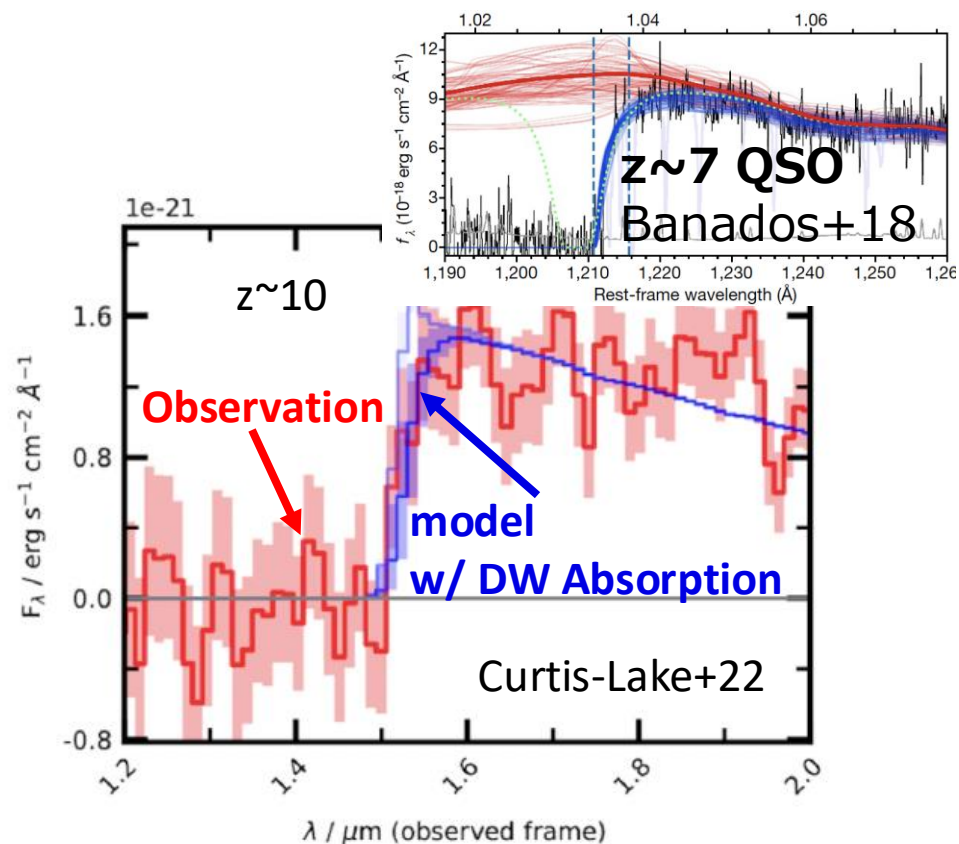
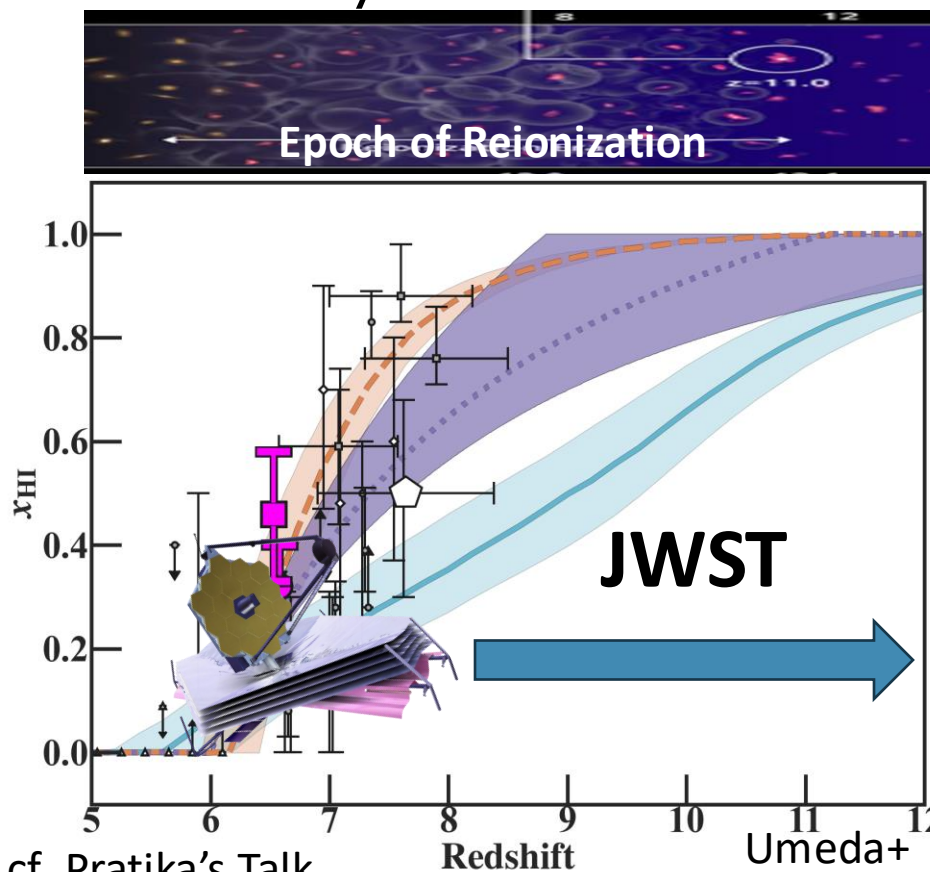
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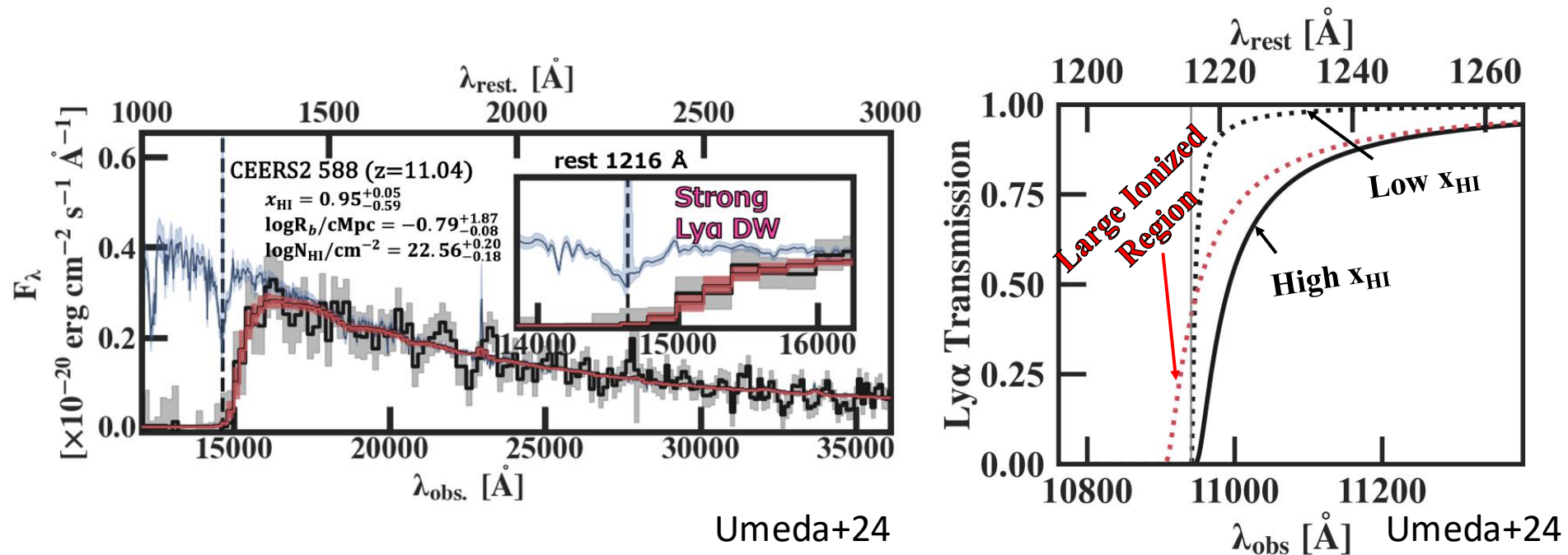


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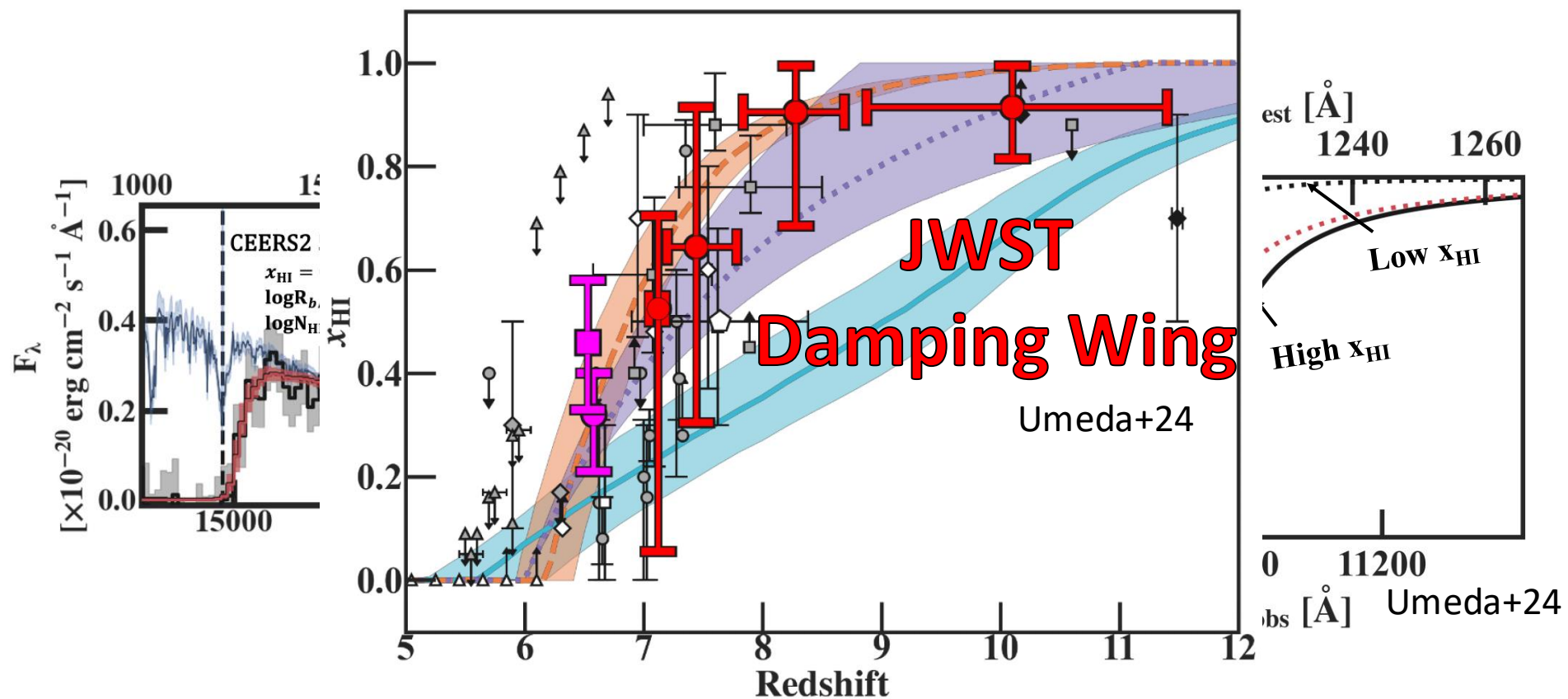
Nature of IGM via Galaxy Observation
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Galaxy UV Continua Absorption As A New Tool

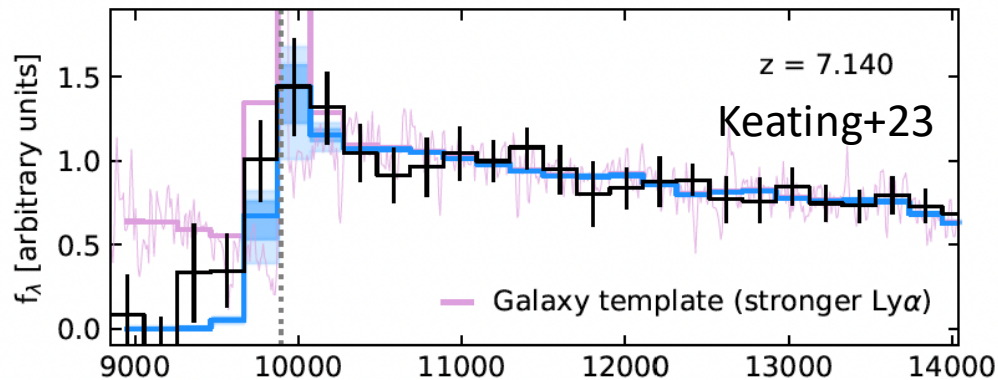


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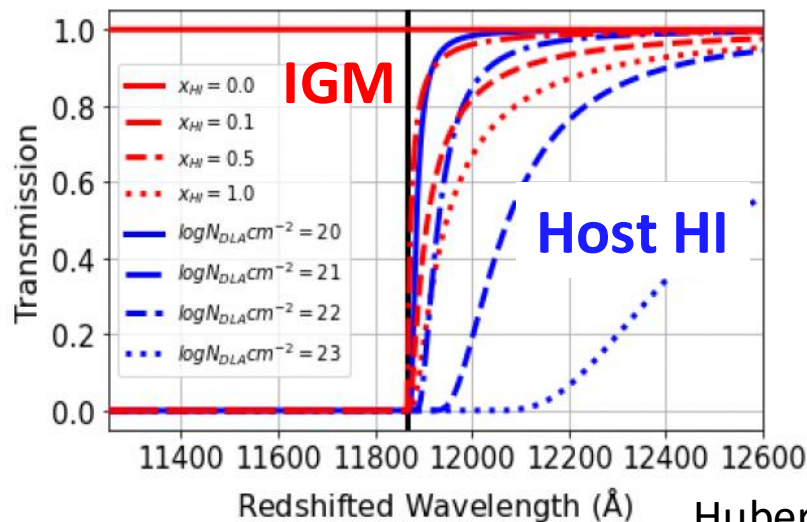


Advances in UV-Continuum Absorption Analysis

Ly α Emission



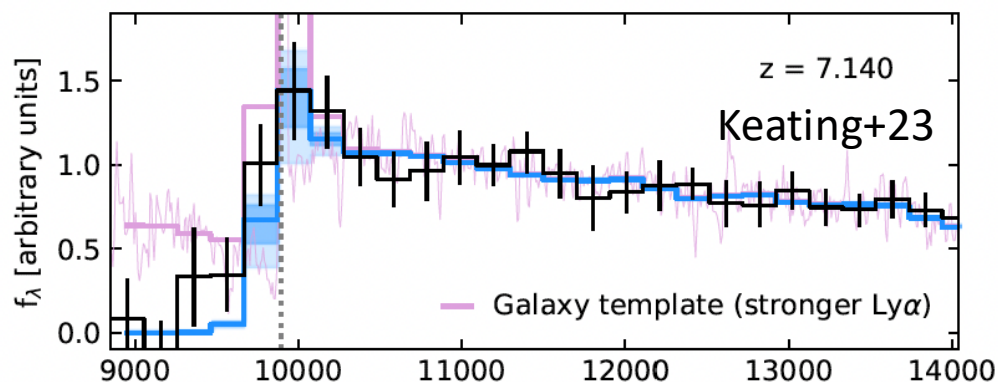
Host Galaxy Absorption



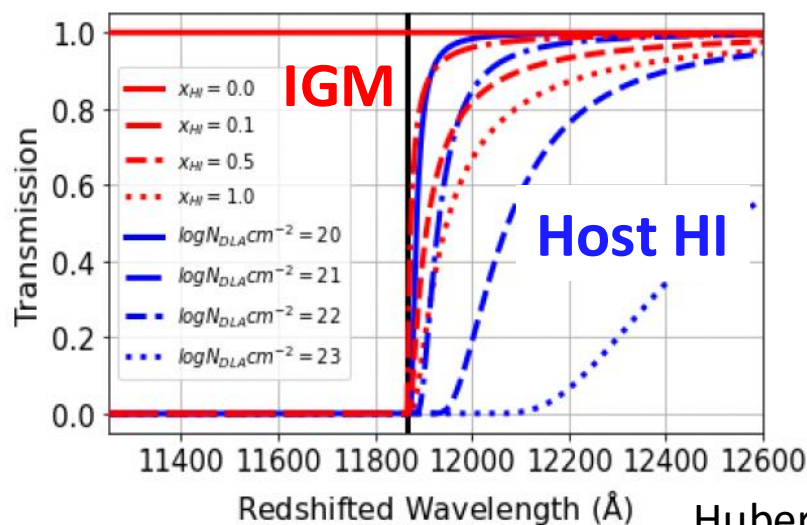
Huberty+25 (cf. Heintz+24, etc...)

Advances in UV-Continuum Absorption Analysis

Ly α Emission



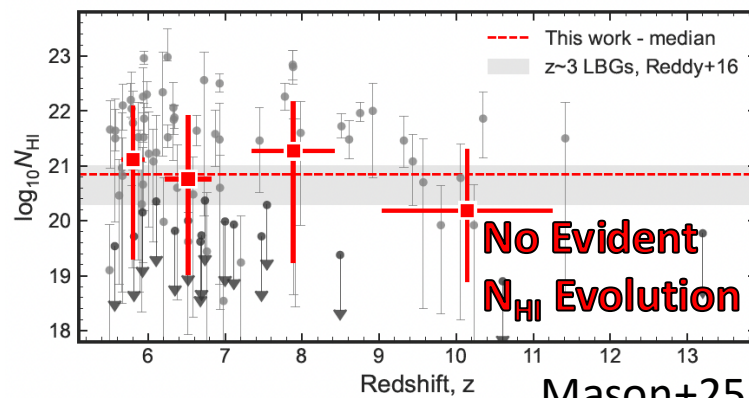
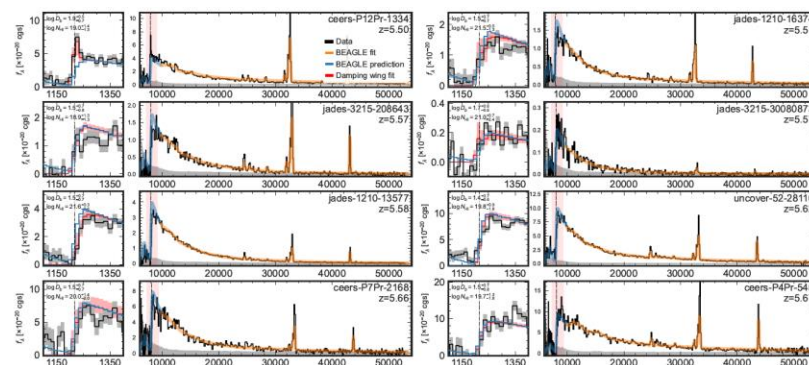
Host Galaxy Absorption



Huberty+25 (cf. Heintz+24, etc...)

Updated Modeling

Bayesian Inference Framework and Semi-Numerical Simulation



Mason+25

(cf. Umeda+24)

Our Strategy: Template-base Analysis

JWST Galaxy Spectra

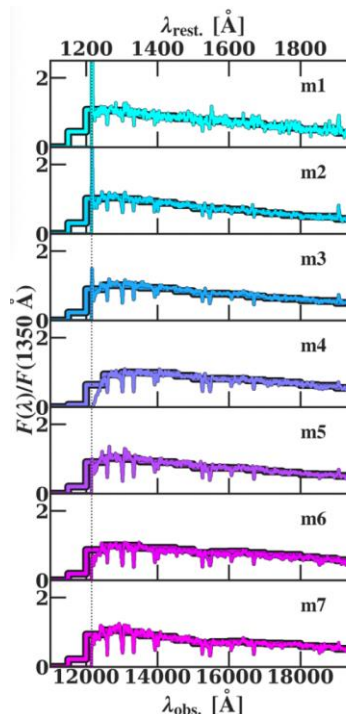
>100 Galaxies at $z > 5$



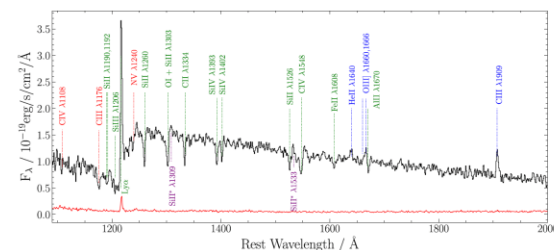
Composite Spectra with
Homogeneous
Subsample
(e.g., Redshift,
UV Magnitude & Slopes)

Modeling

Realistic Template w/
Host HI Absorption



Composite Spectra from
>600 Galaxies at $z=2-5$
(VANDELS; Cullen+19)



Data & Sample

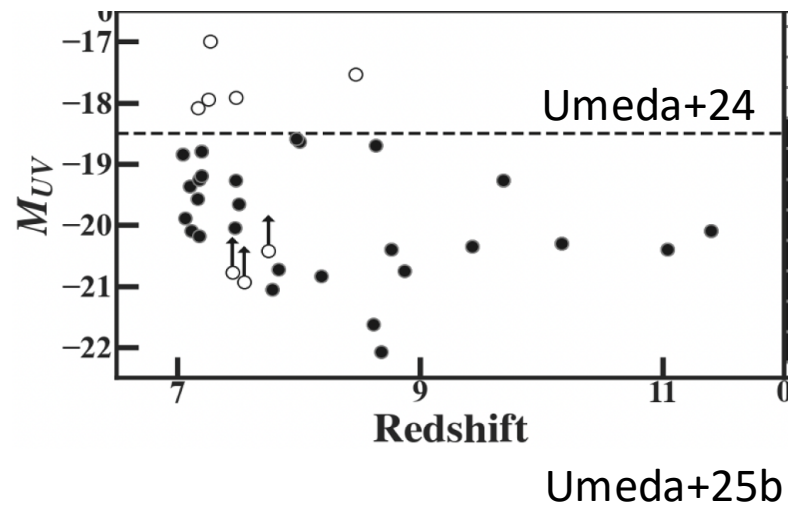
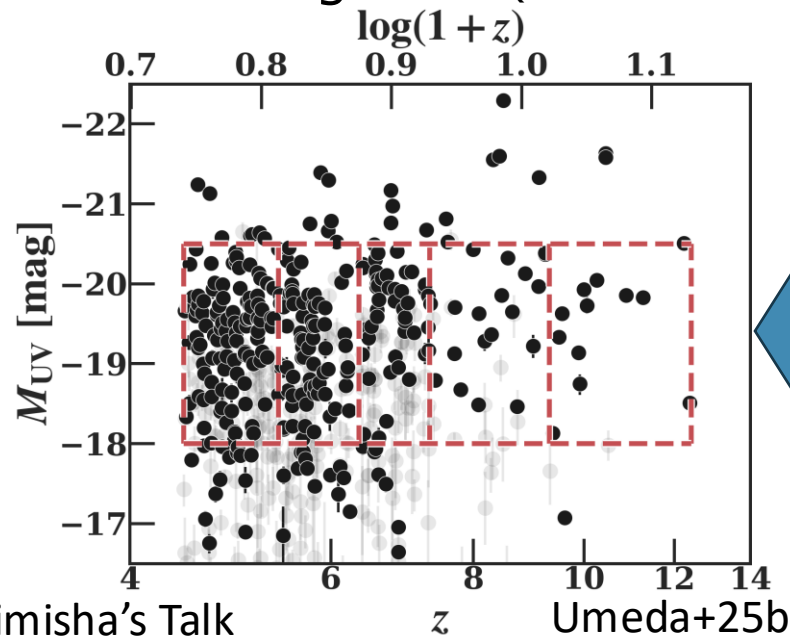
PRISM & Spec-z by Emission $\rightarrow$ 581 Objects in $4.5 < z < 13$ Umeda+25b

Thx to JADES, UNCOVER, CEERS, DD, GO Teams

UV Magnitude/Slope Measurement (Method Similar to Saxena+24)

Fiducial Subsamples: $-20.5 < M_{UV} < -18$, $-1.8 < \beta < -2.6$

Omit Strong LAEs (Measured in same way as Kageura+25)



Data & Sample

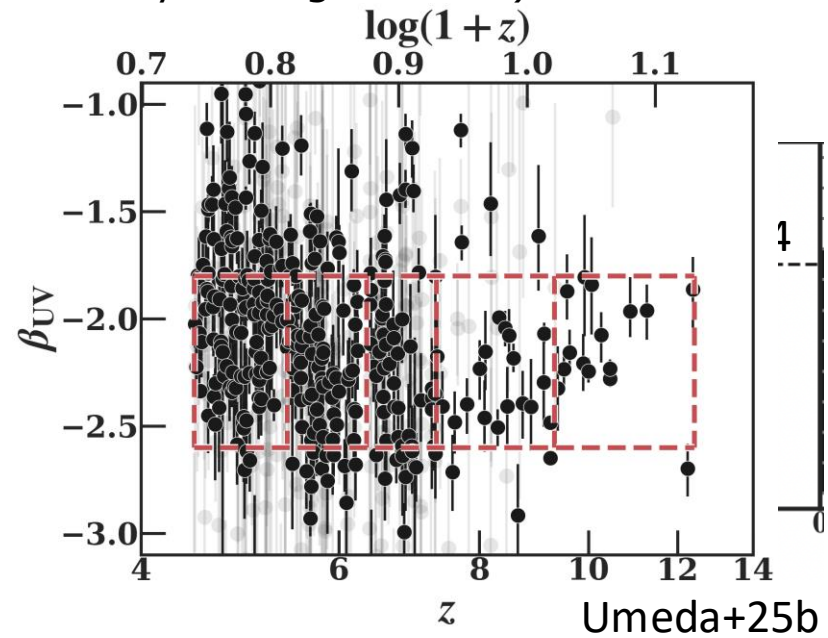
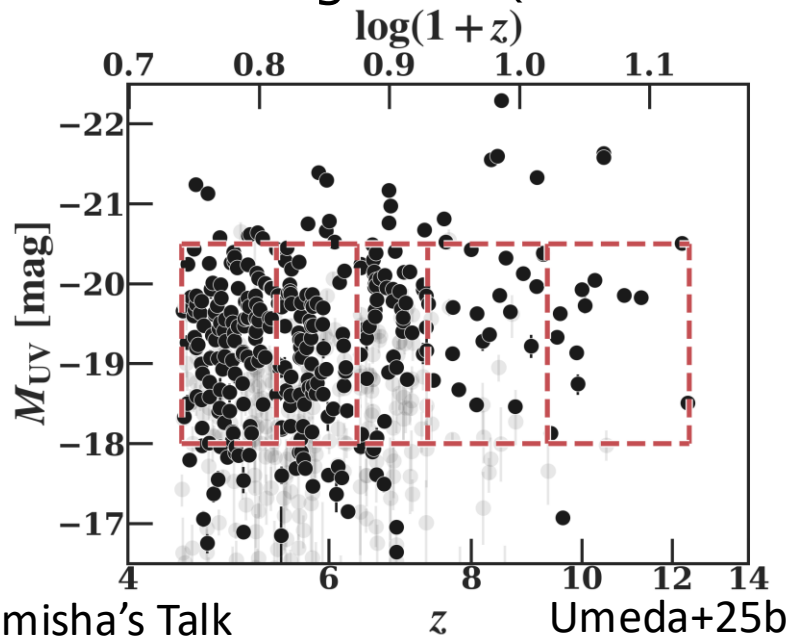
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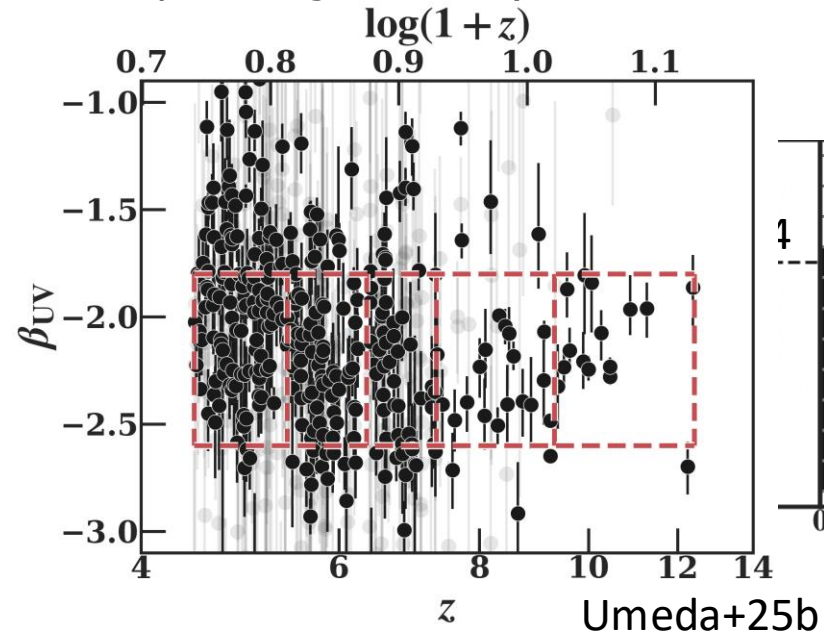
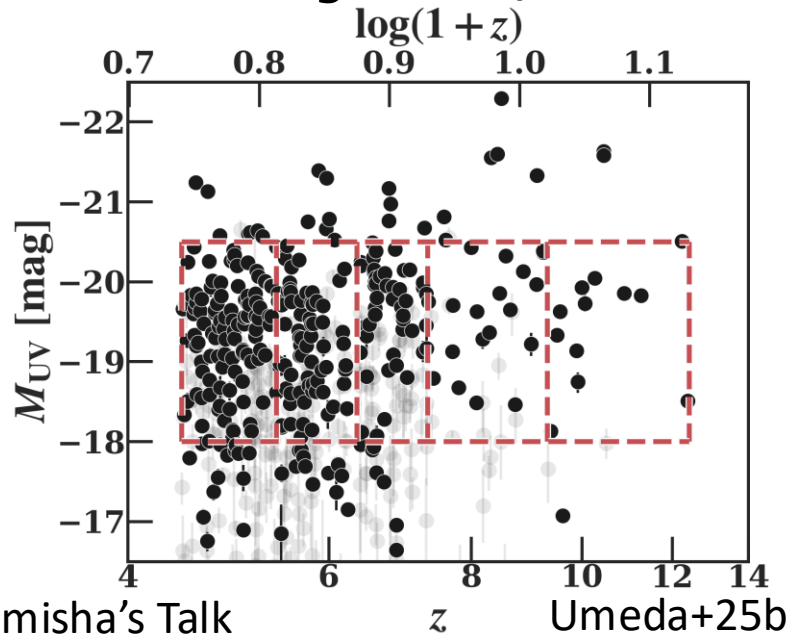
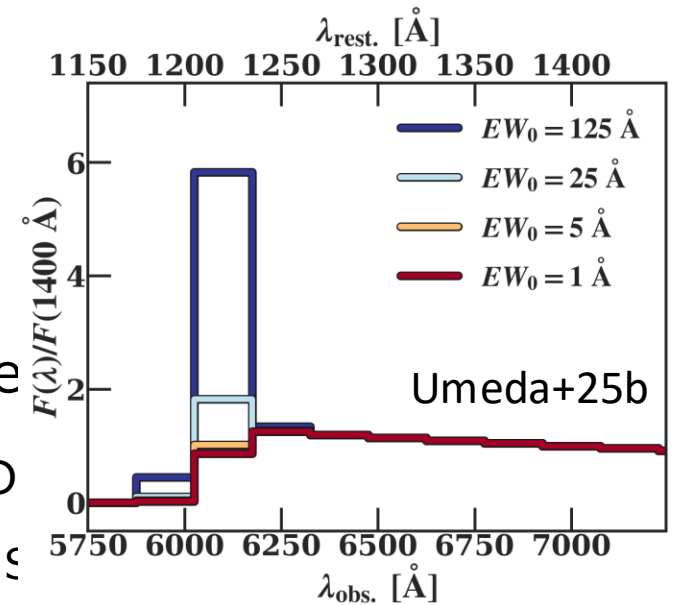
PRISM & Spec-z by Emission → 581 Objects

Thx to JADES, UNCOVER, CEERS, D

UV Magnitude/Slope Measurement (Method 5)

Fiducial Subsamples: $-20.5 < M_{UV} < -18$, $-1.8 < \beta < -2.6$

Omit Strong LAEs (Measured in same way as Kageura+25)



Data & Sample

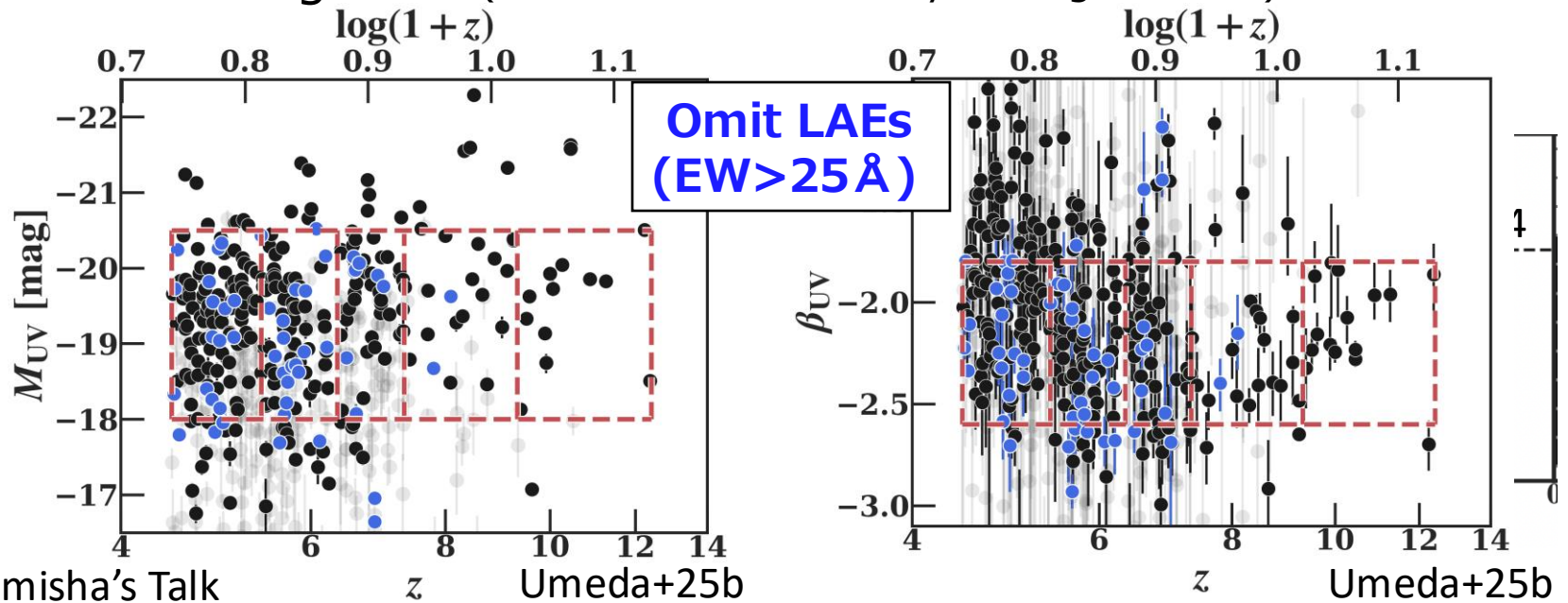
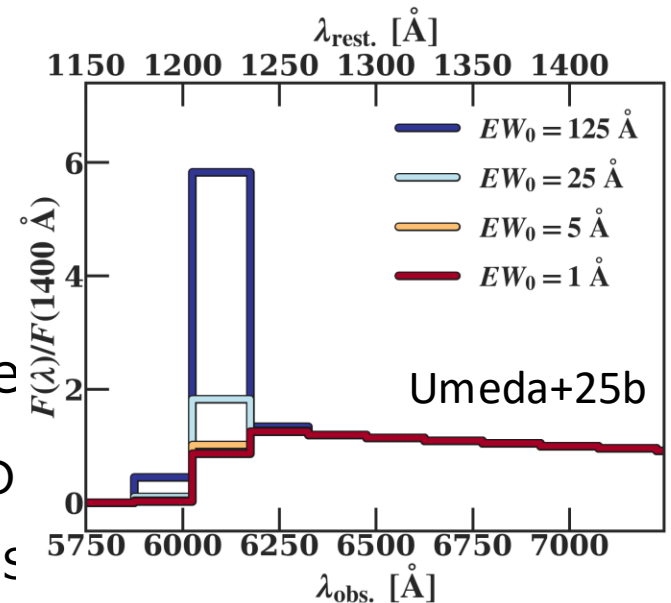
PRISM & Spec-z by Emission → 581 Objects

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UV Magnitude/Slope Measurement (Method 5)

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Omit Strong LAEs (Measured in same way as Kageura+25)



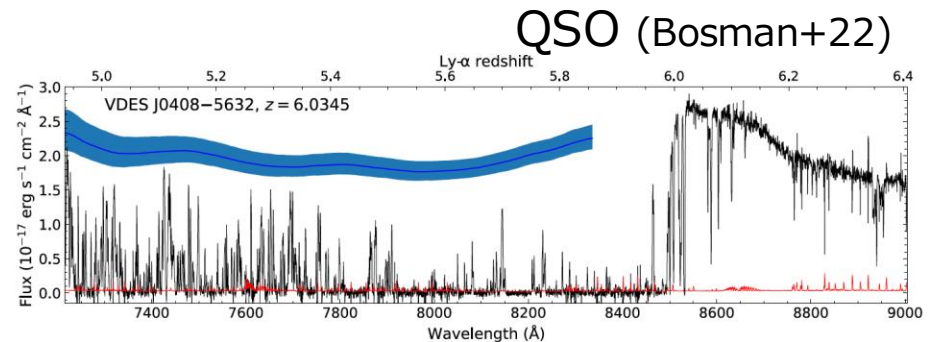
cf. Nimisha's Talk

“Tail” of Reionization Ly α Forest Signals

Bright Looking As Quasar→Ly α Forest!

Ly α Forest Signals with stack data ($\Delta z=0.5$, $4.5 < z < 6.0$)

Dissipating Signals→Witnessing the End of EoR

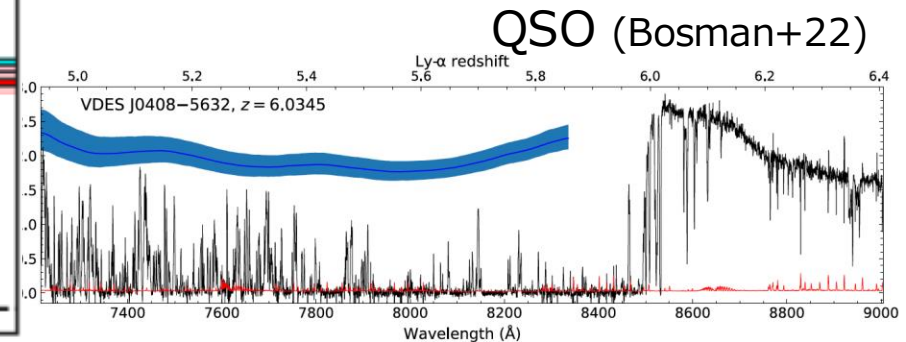
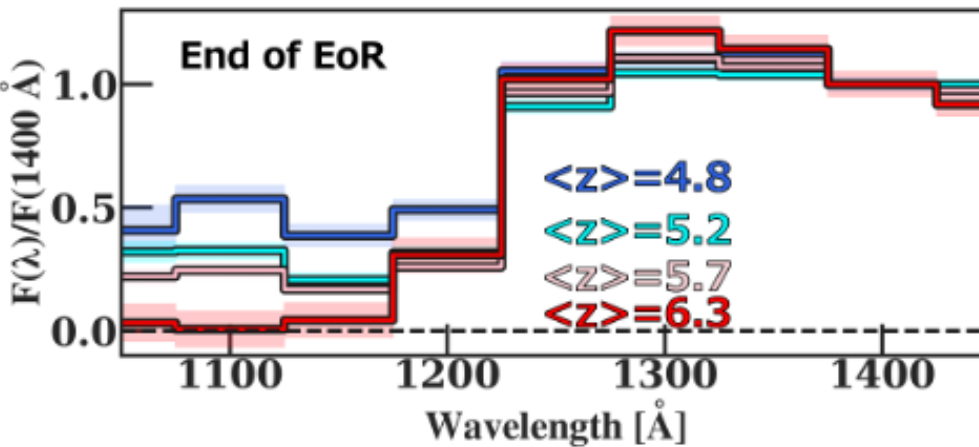


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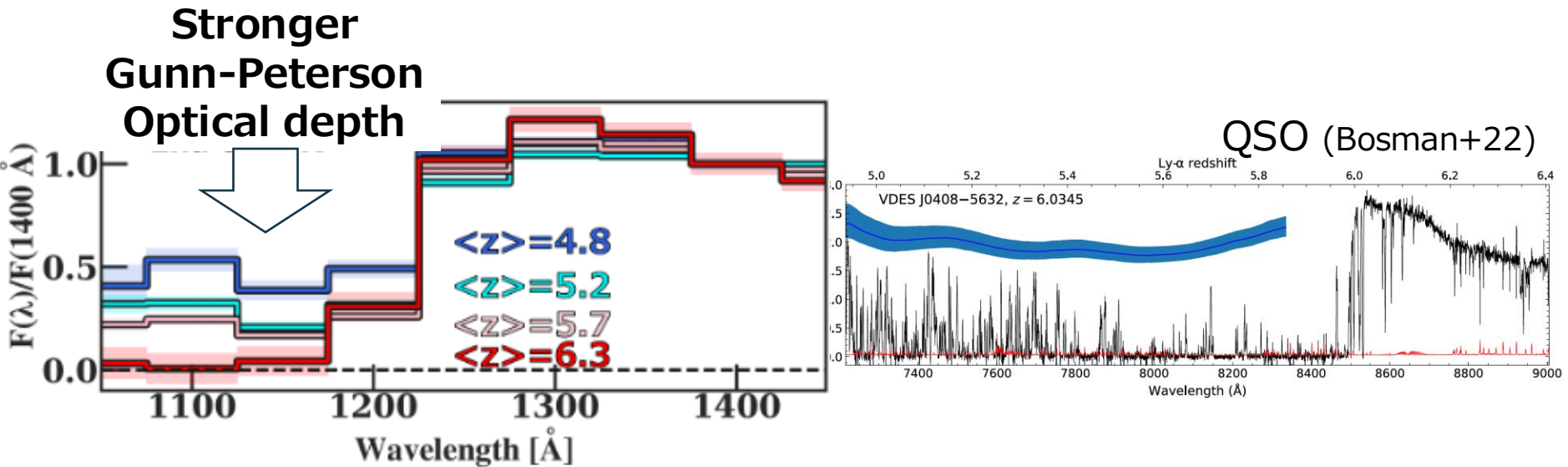
Umeda+25b

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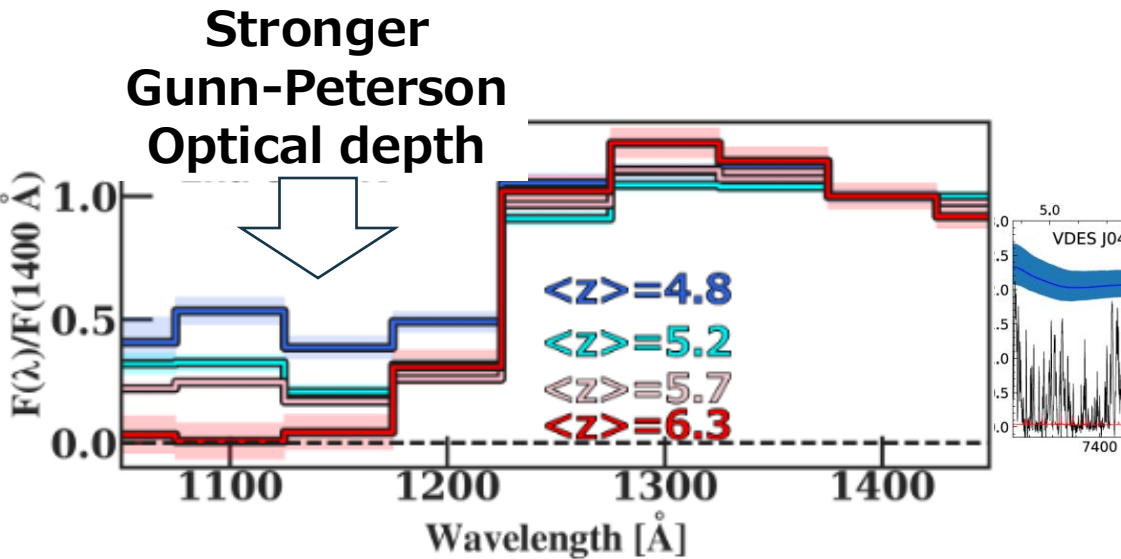
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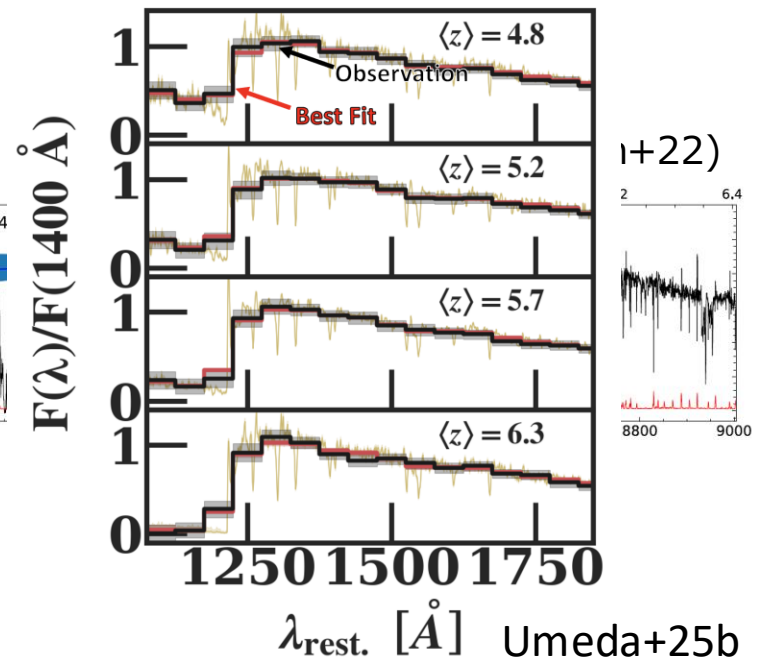
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Umeda+25b



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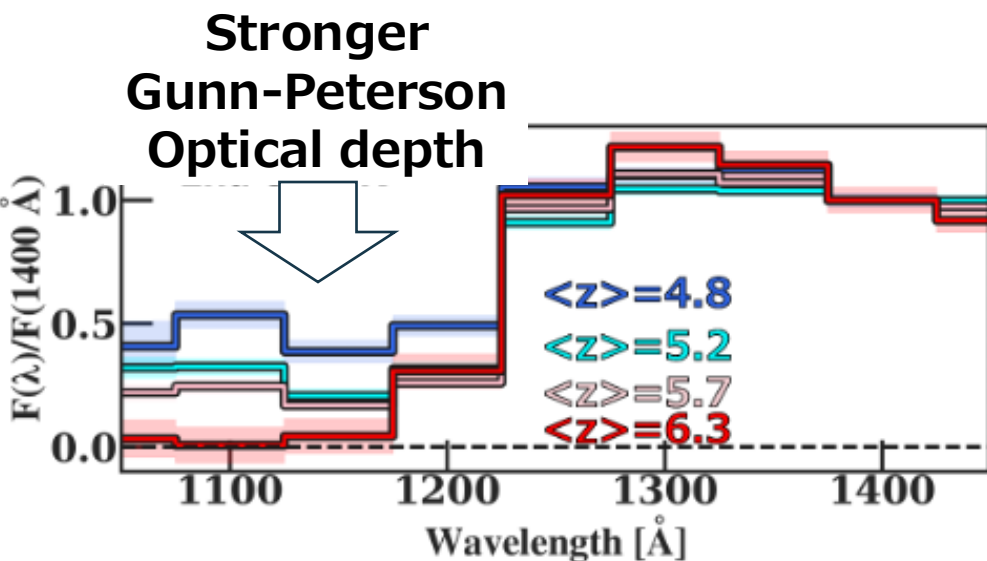
(cf. Meyer+25, Mason+25)

“Tail” of Reionization Ly α Forest Signals

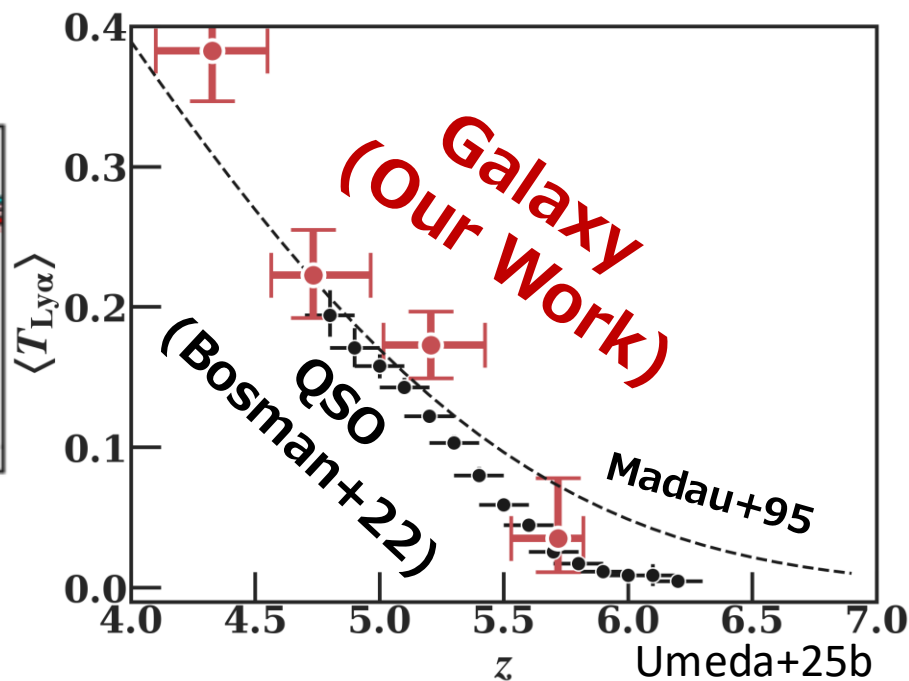
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Umeda+25b



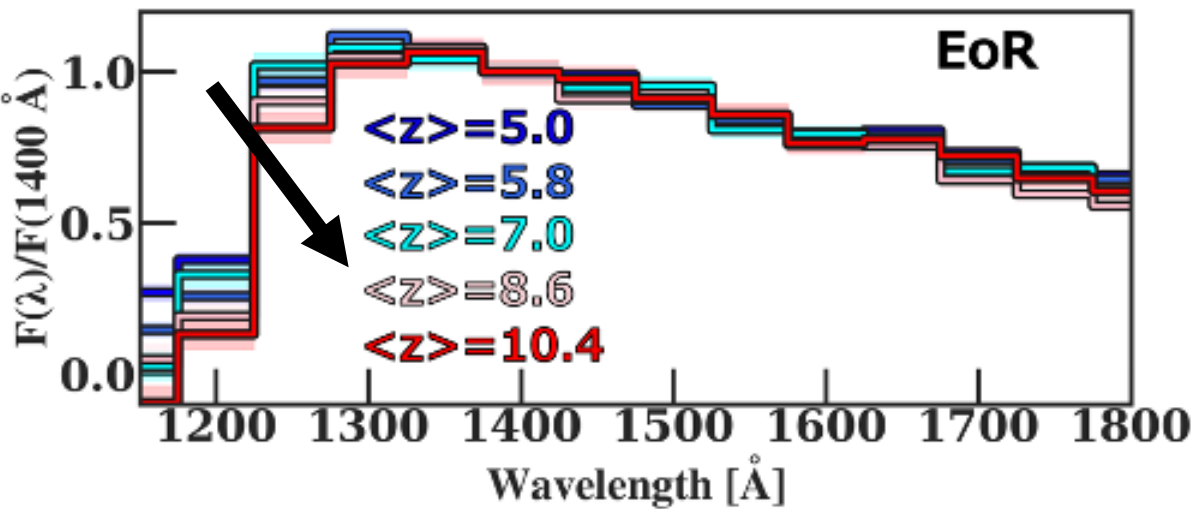
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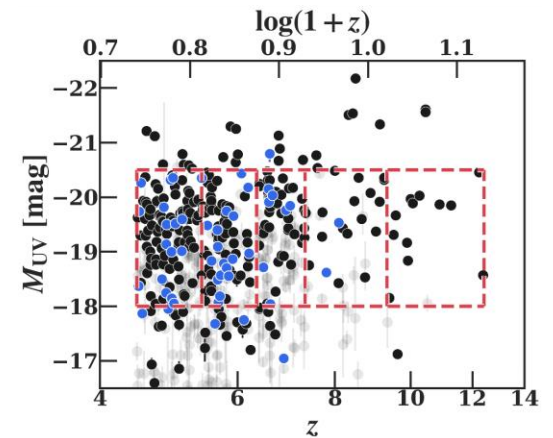
Deep Into EoR

Lya Break Evolution

Lyman Break Softening
Toward High- z



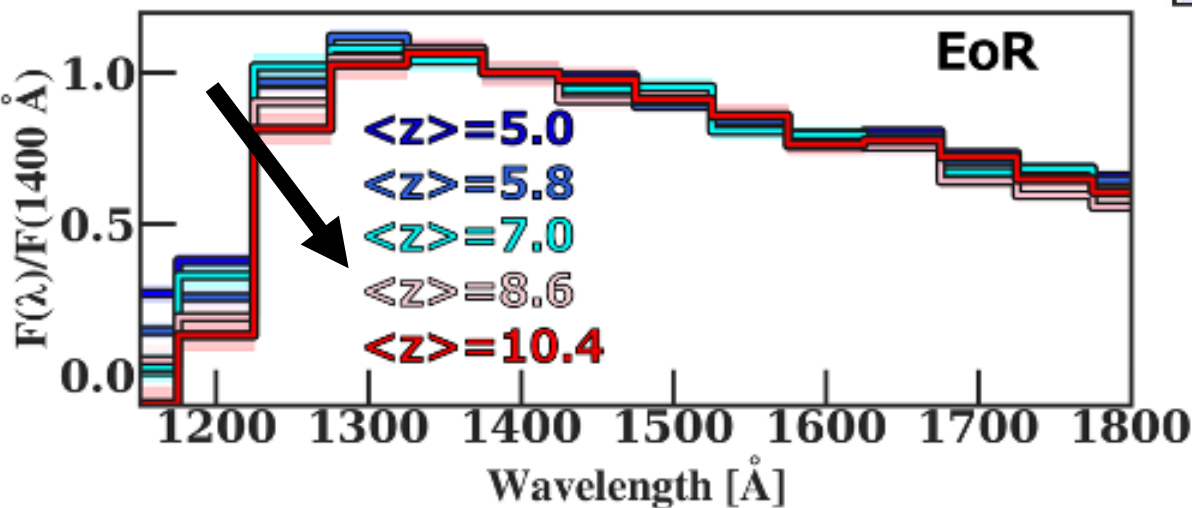
Umeda+25b



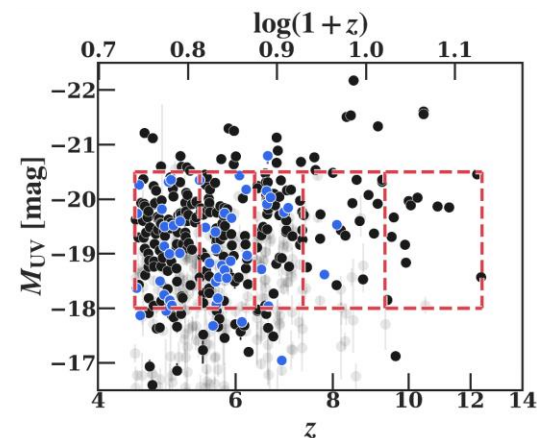
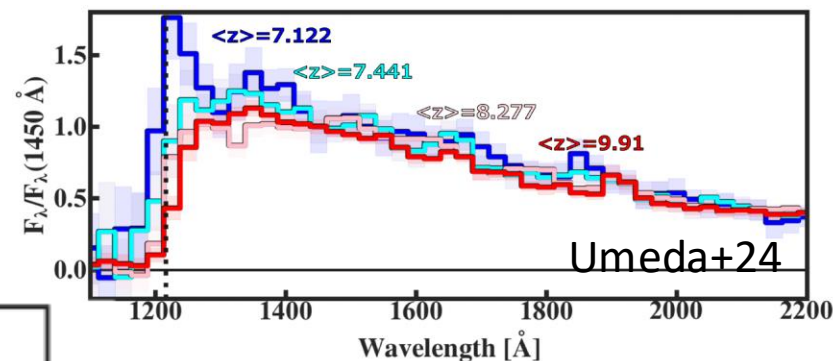
Umeda+25b

Deep Into EoR *Ly α Break Evolution*

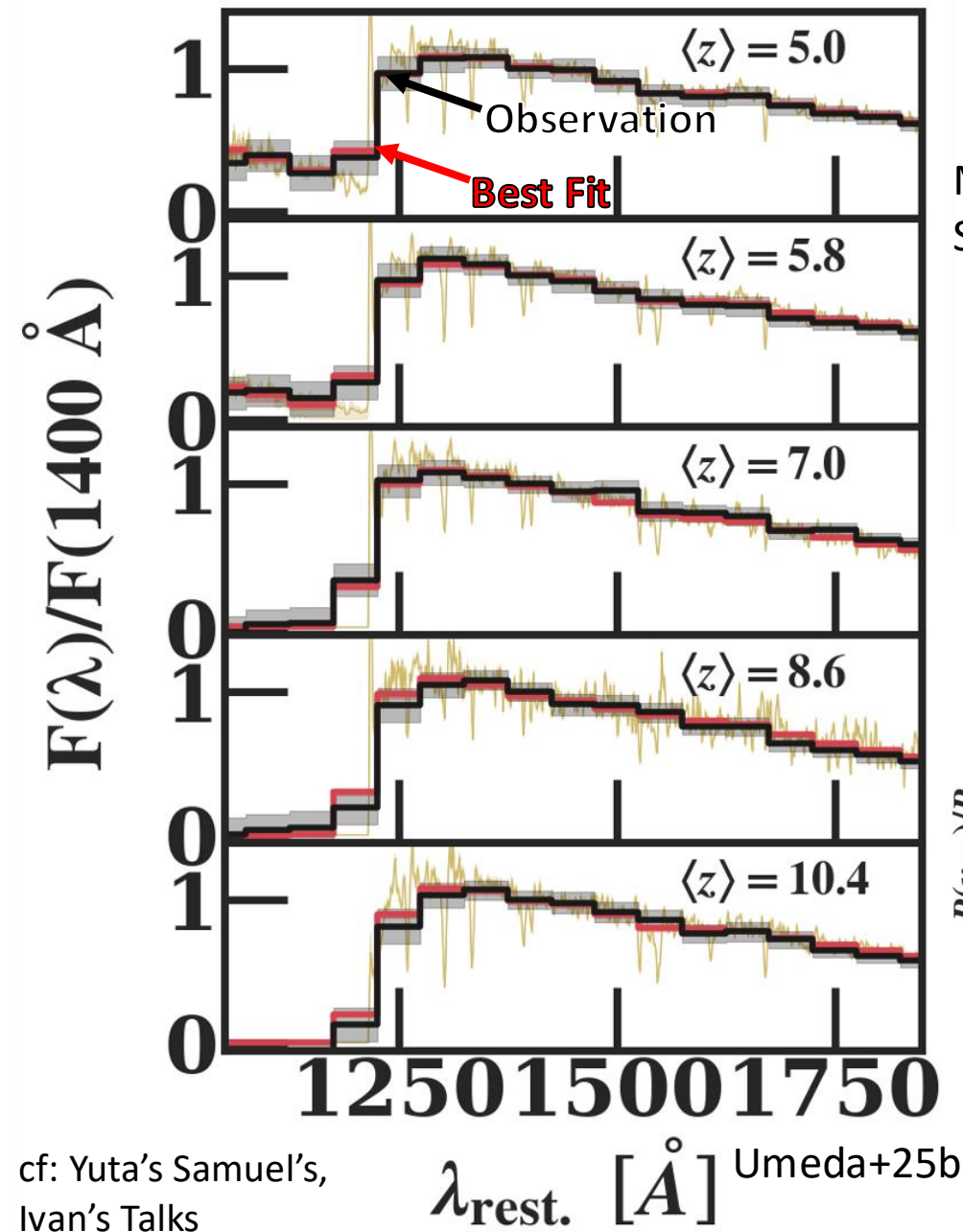
Lyman Break Softening
Toward High- z



Umeda+25b



Umeda+25b



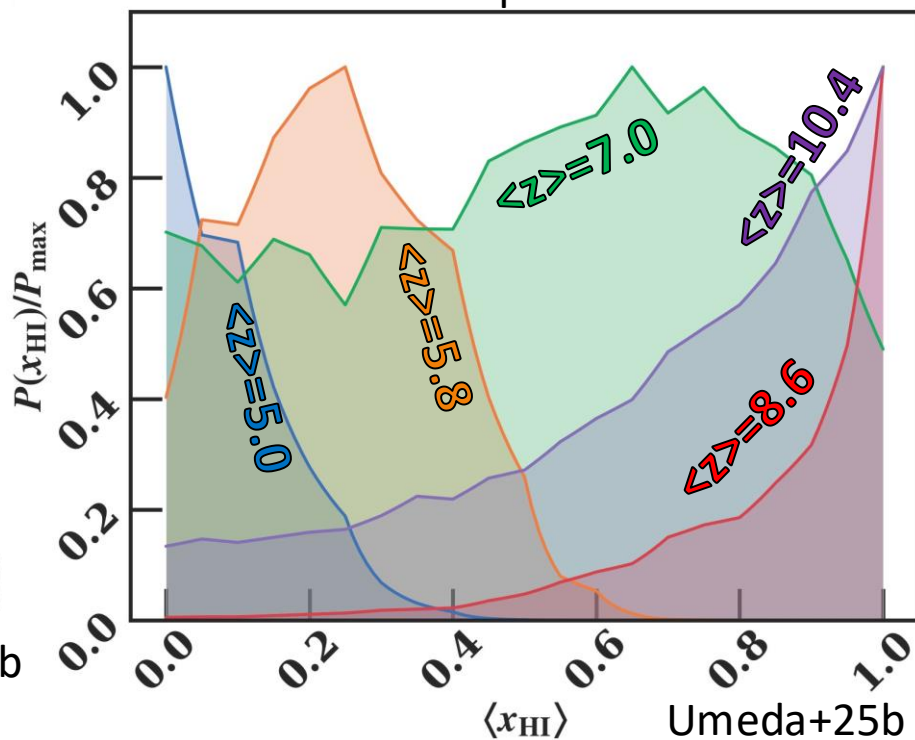
x_{HI} Inference

Mock Observations (cf. Yuta's Talk):

Sightline by Muv→Stack

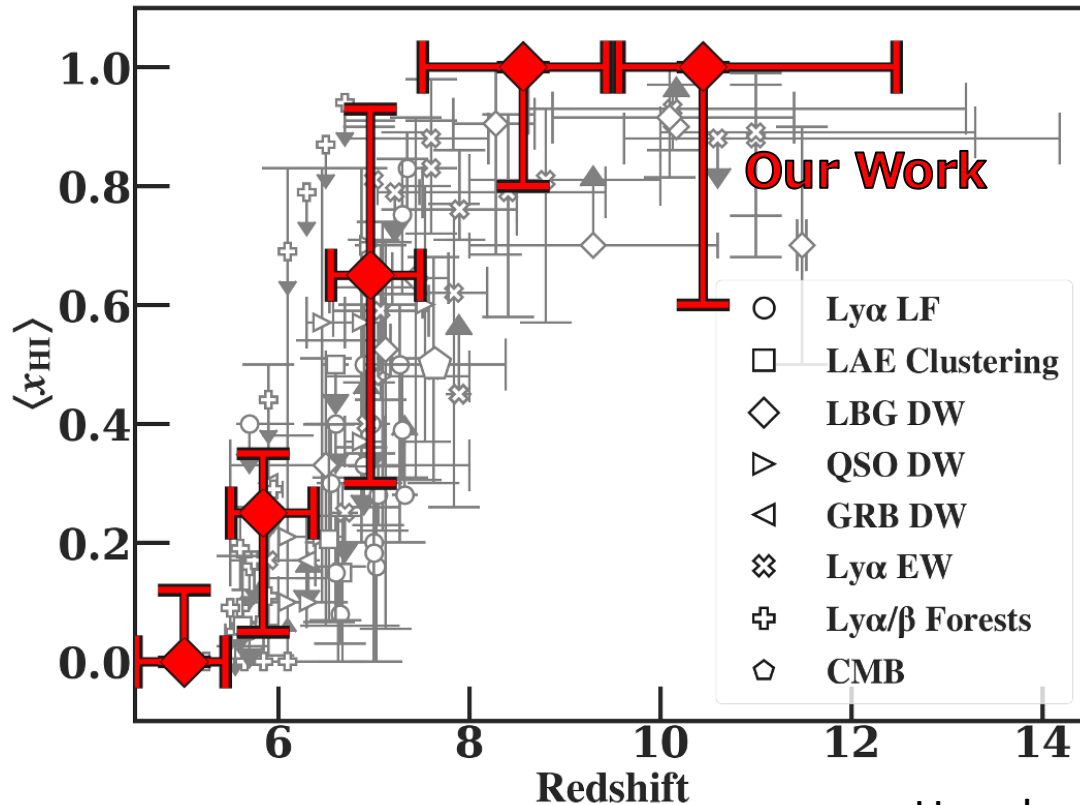
IGM Model: 21cmFAST (Murray+20)

- Template Spectra: $z \sim 2-5$ Galaxies (VANDELS; Cullen+19)
- Omit LAE-like sightlines (Mason+18)
- Residual HI Absorption Considered



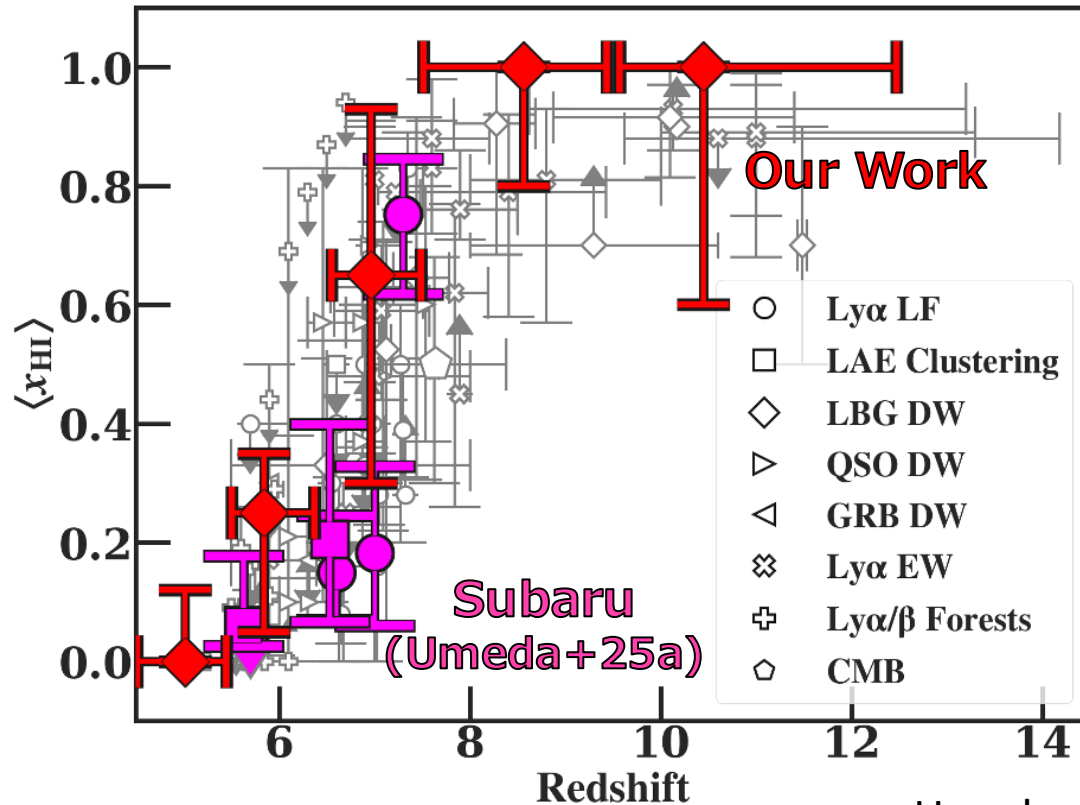
Overviewing Entire Cosmic Reionization History

✓ JWST Ly α Emission (Nakane+24, Kageura+25, Tang+24ab, Napolitano+24, Jones+24,25)



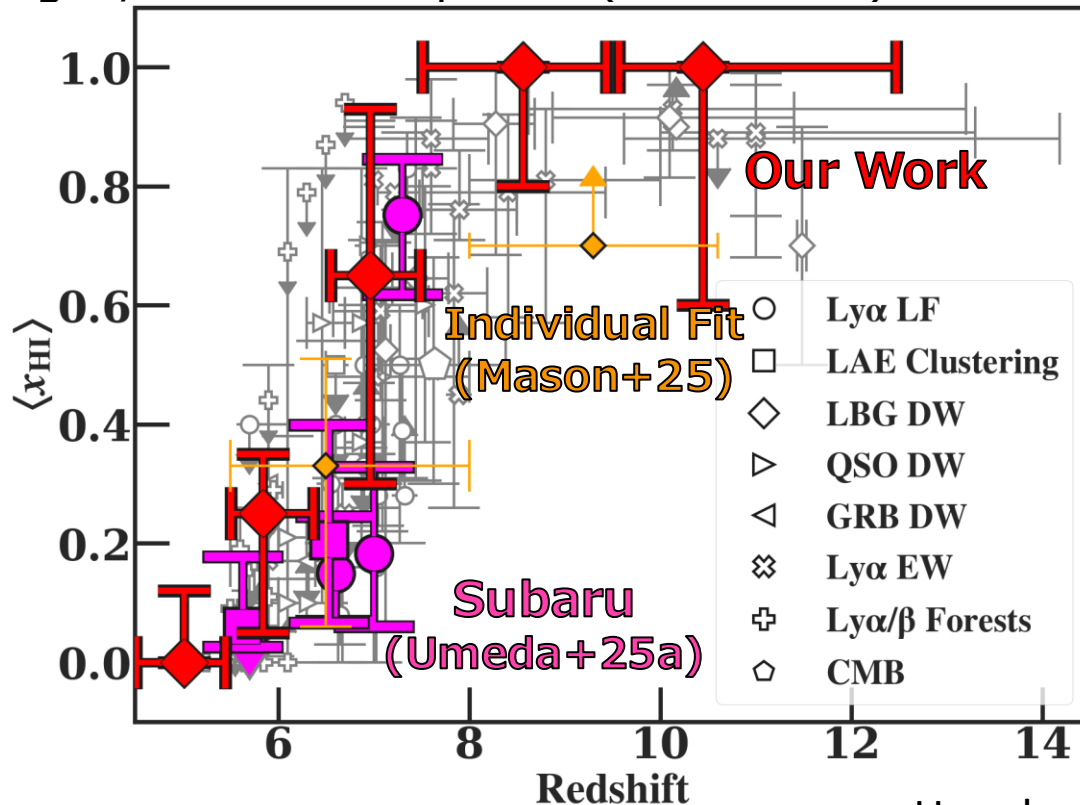
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- ✓ Subaru Wide-Field LAE Survey (Umeda+25a)



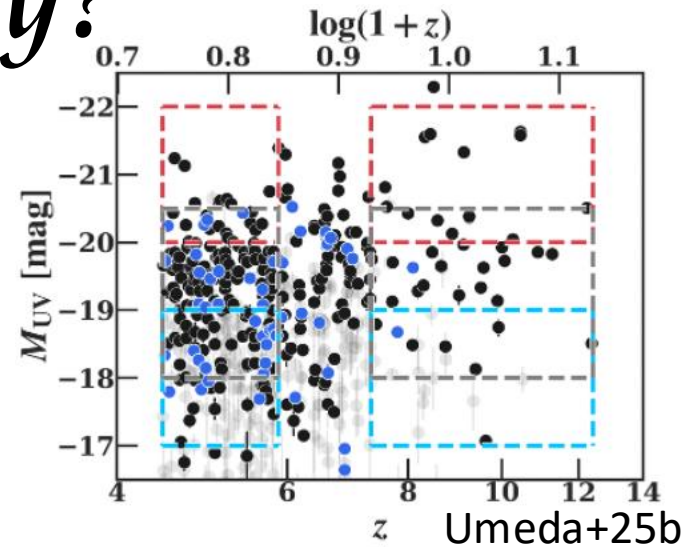
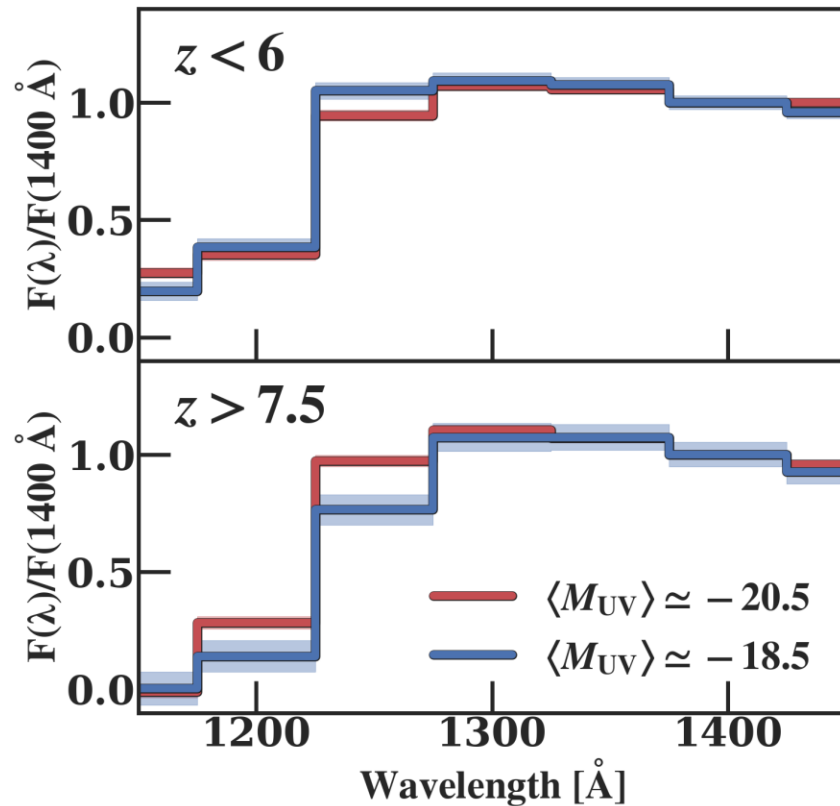
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- ✓ Subaru Wide-Field LAE Survey (Umeda+25a)
- ✓ DW Fitting w/ Individual Spectra (Mason+25)



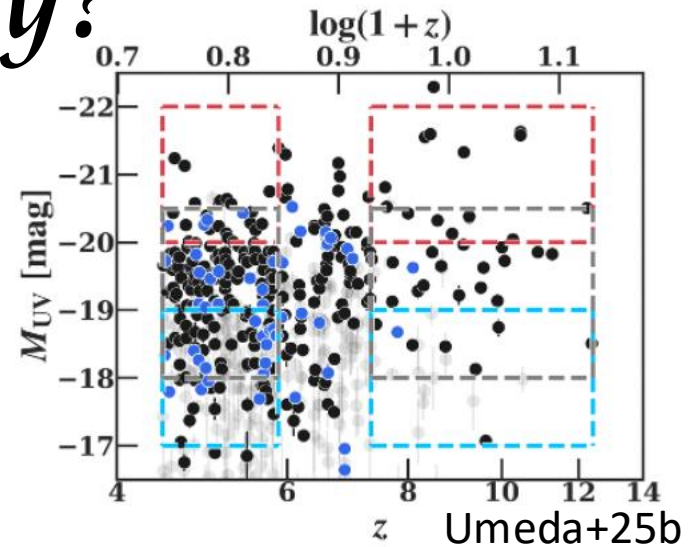
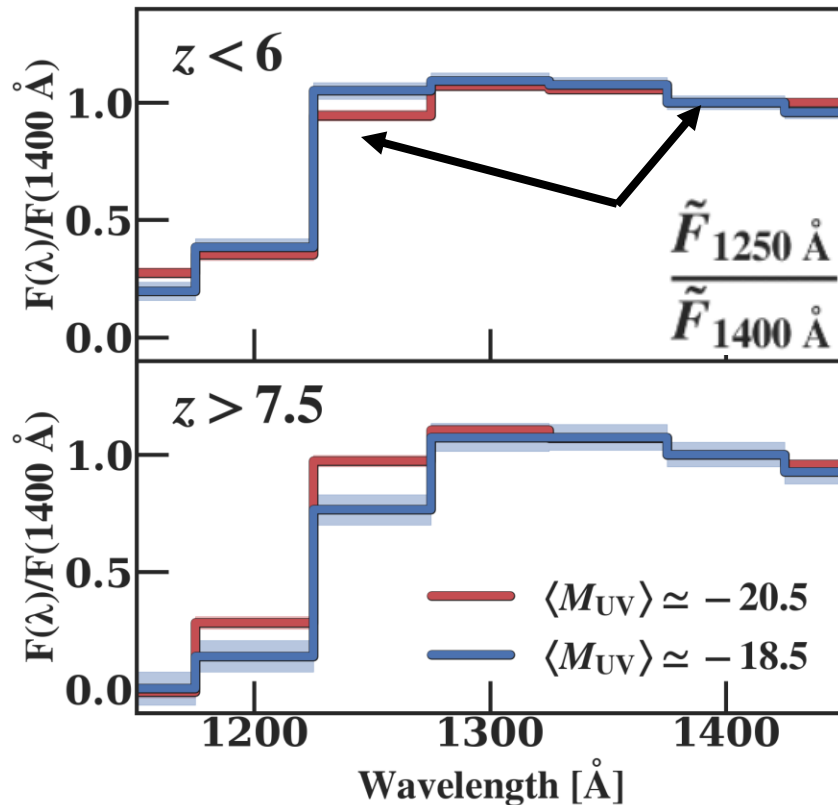
Topology?

Ionization Around Light/Massive Halo
→ Topology of Ionized IGM



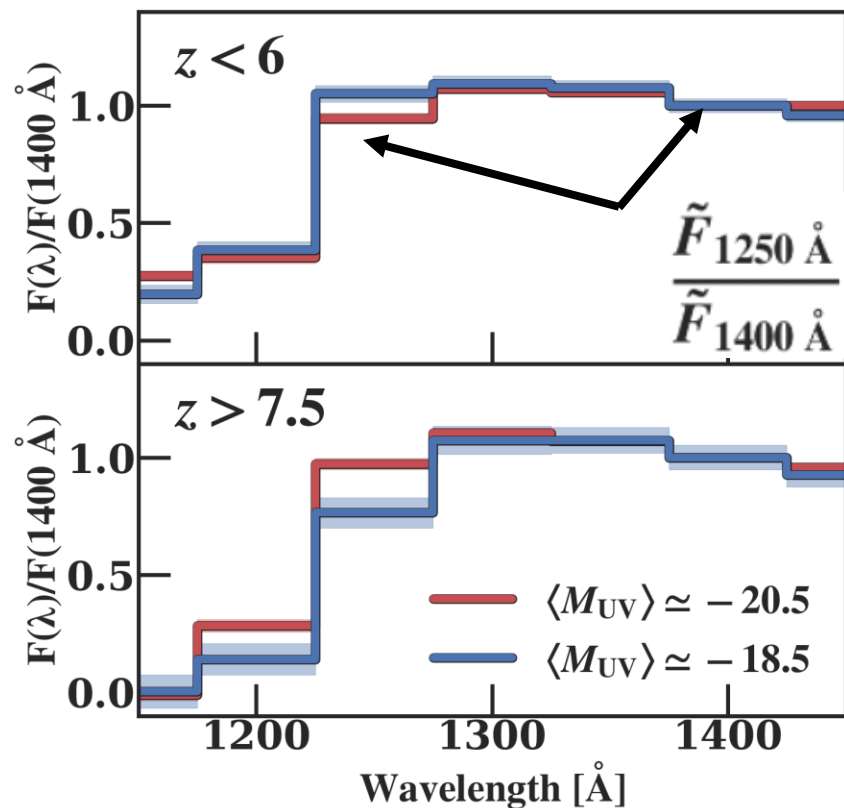
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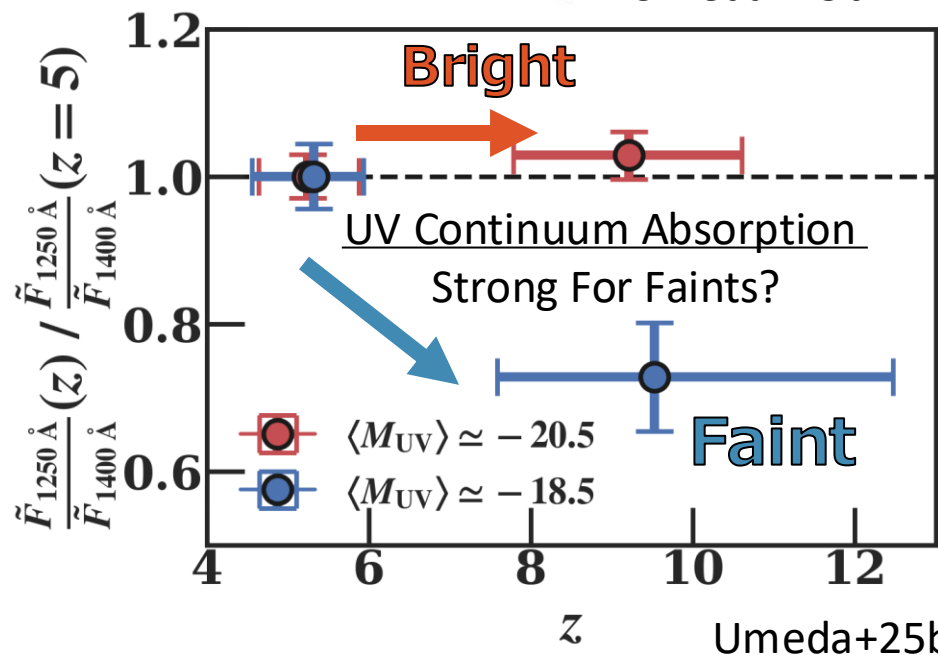
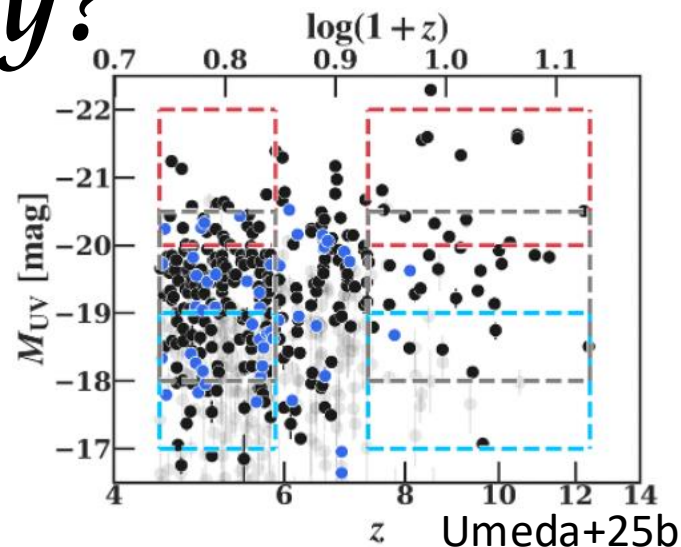
Topology?

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cf. Emma's Talk

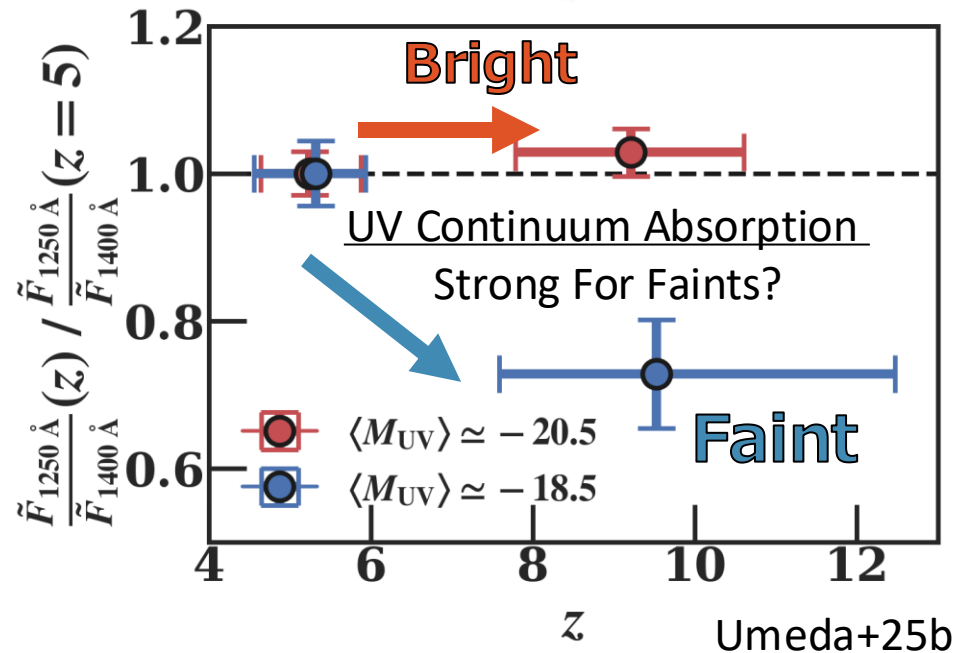
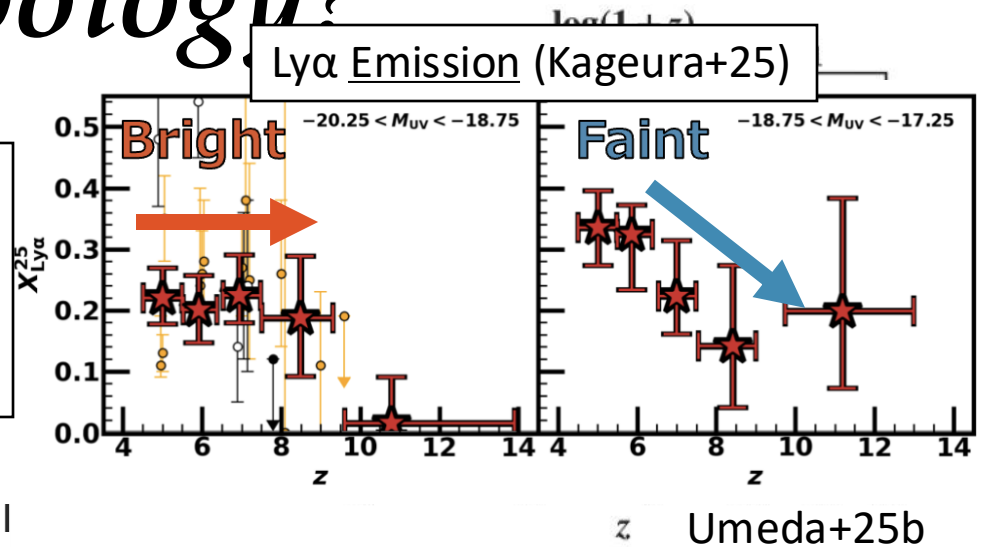
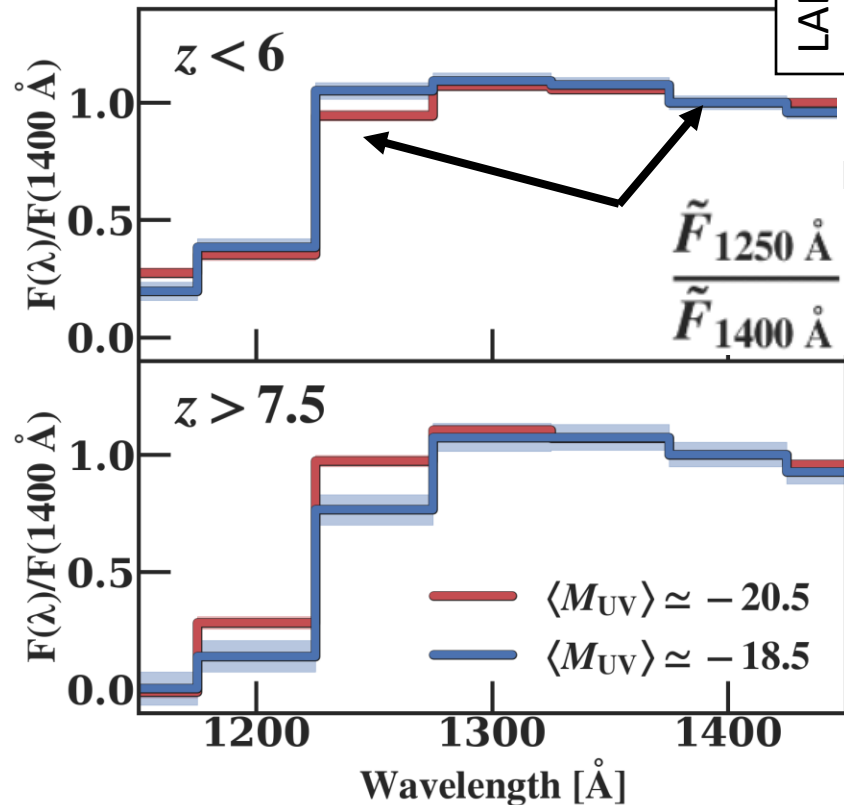
Umeda+25b



Umeda+25b

Topology?

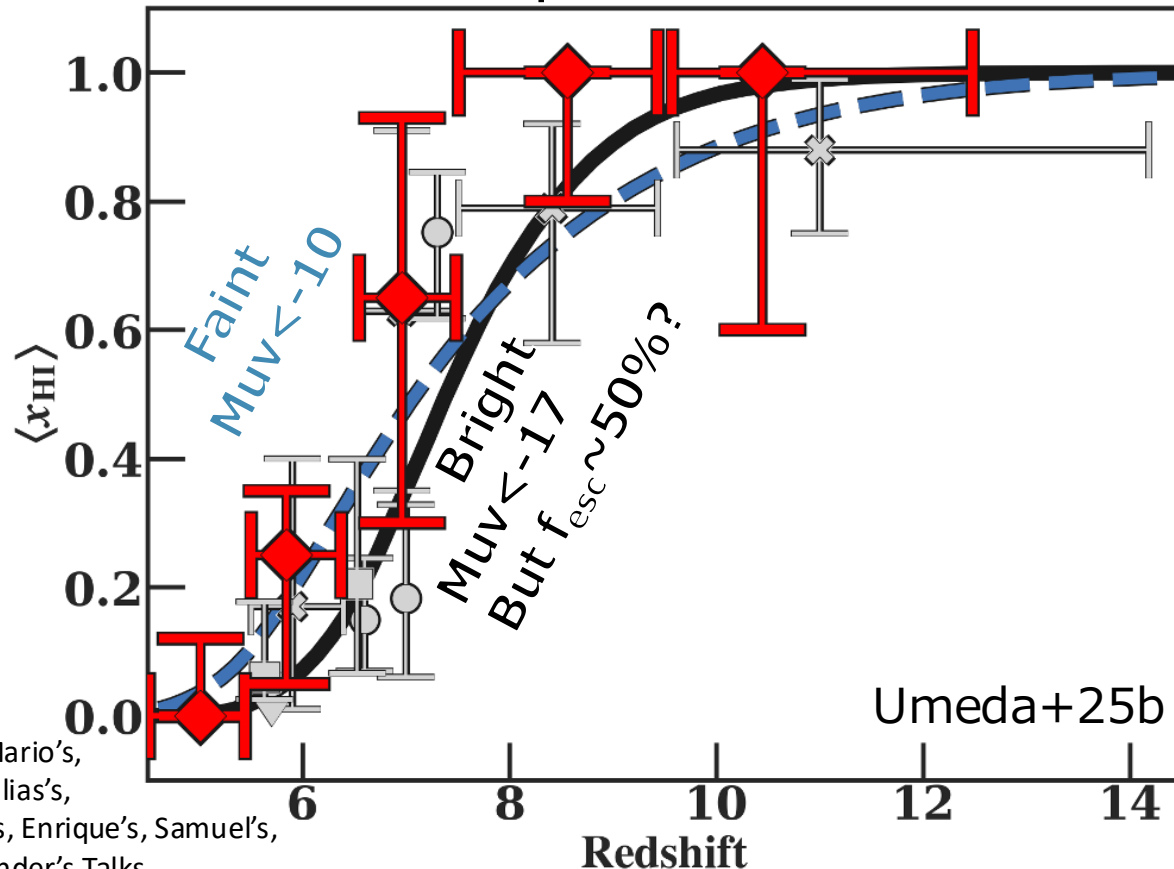
Ionization Around Light/Massive Halo
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Cosmic Reionization Too Rapid and Too Late?

Continuum Absorption, Ly α EW, Ly α LF/Clustering

$\Rightarrow$ All Favors Too Rapid & Late Reionization?

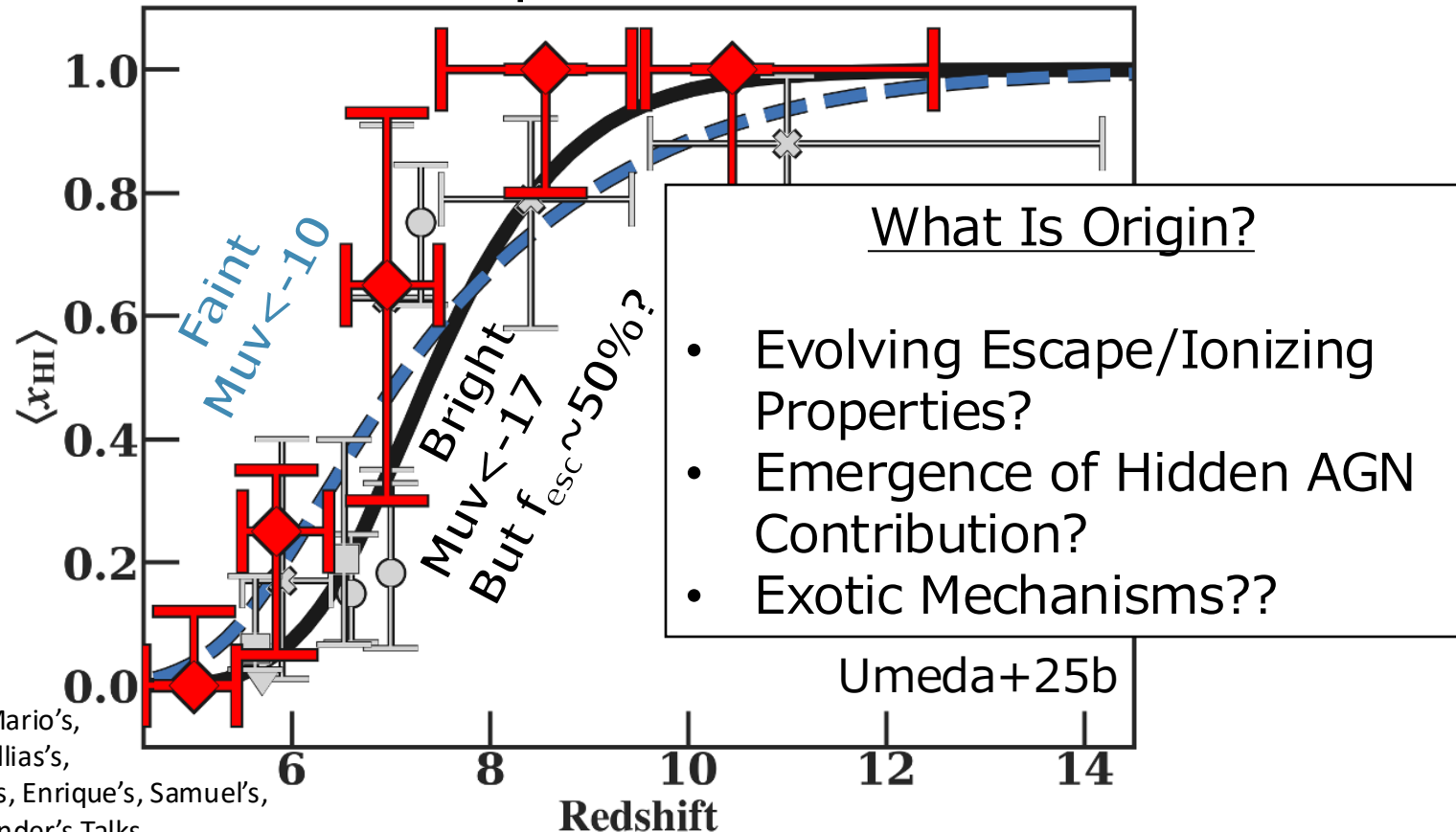


cf. Andrea's, Charlotte's, Mario's,
Eros's, Karla's, Valentin's, Illias's,
Seonwoo's, Emma's, Juan's, Enrique's, Samuel's,
Yuta's, Alessandra's, Alexander's Talks

Cosmic Reionization Too Rapid and Too Late?

Continuum Absorption, Ly α EW, Ly α LF/Clustering

$\Rightarrow$ All Favors Too Rapid & Late Reionization?



Αποχωρώντας από την ομιλία

Up On arXiv This Week!

$z < 6$: Gunn-Peterson Trough Strengthen

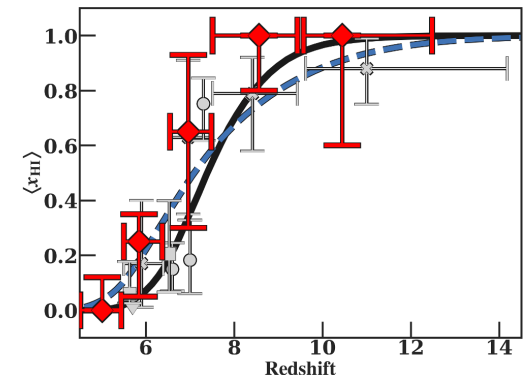
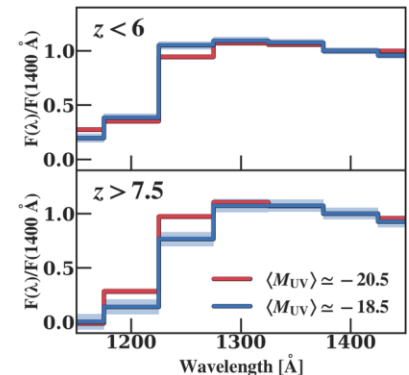
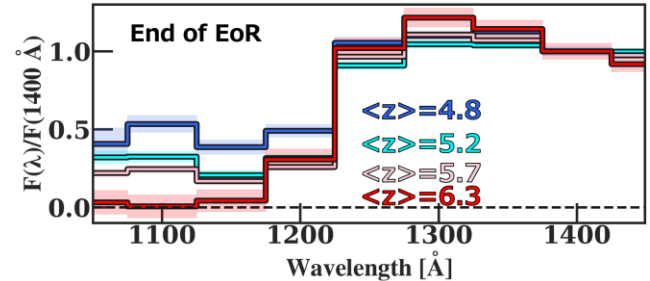
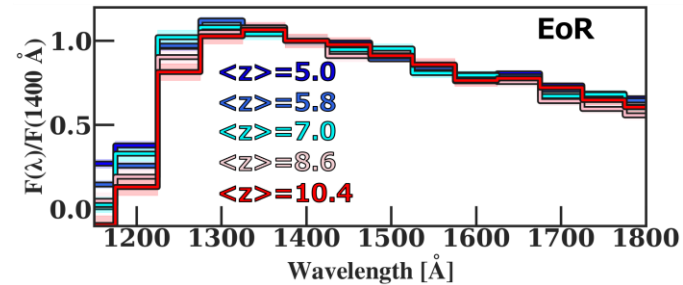
$z > 7$: Ly α Damping Wing Strengthen

Dependencies on Galaxy Properties
(UV magnitudes)

- Bright Galaxies Dominant Driver of Cosmic Reionization?

Rapid Reionization History Still Supported

- Too Rapid & Late to Reproduce?
- Hidden Ionizing Sources Emerging at $z \sim 8$?
- Evolving Escape Fractions / Ionizing Photon Production Efficiency?
- Any Ideas?



Escaping from Job Hunting Labyrinth

★ FINISHING PHD! Job Market from Next Season! ★

If $\sigma_{\text{interest}} > 0$ AND $\tau_{\text{contract}} \gtrsim 0(1 \text{ yr})$,

Please let me know!

Keywords:

Cosmic Reionization

Subaru

JWST

$\text{Ly}\alpha$ Emitters

High- z ($z > 7$)



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Homepage:

<https://astroumeda.github.io>

